The Y-system for Scattering Amplitudes

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Planar gluon scattering amplitudes (MHV) are given by Null Polygonal Wilson Loops

[Alday,Maldacena; Drummond,Henn,Korchemsky,Sokatchev; Berkovitz,Maldacena; Beisert,Ricci,Tseytlin,Wolf; ...]



Classical Integrability at strong coupling

Flat connection

Spectrum

 $\Omega(x)$

$$A = \frac{J + *J}{x - 1} + \frac{J - *J}{x + 1} , \qquad J = -g^{-1}dg$$

Amplitudes



Summary of results

$$Area_{reg} = \sum_{k} \int d\theta \ m_k \cosh \theta \log(1 + Y_k(\theta))$$

 $[\text{Cross ratio}]_s = Y_s(0)$

$$\log Y_k(\theta) = -m_k \cosh \theta + C_k + K_{k,k'} \star \log(1 + Y_{k'})$$

Simple integral equation for a set of 3n-15 generalized cross ratios $Y_k(\theta)$ in terms of 3n-15 parameters m_k , C_k

Takes the form of **Thermodynamic Bethe Ansatz equations** for an integrable model with **relativistic** particles of masses given by the m's. The Area is the corresponding **Free Energy!**

Very suggestive....



 $e^{2\theta} = \zeta^2 = \frac{x-1}{x+1}$

Example (in $R^{1,1}$ kinematics) Parametric solution (10 gluons) 0.6 0.6 0.5 2.0 0.5 A 0.4 A 0.4 0.3 0.3 1.0 0.2 0.2 .0 m_2 0.5 $Y_{2}(0)$ 0.5 1.0 0.5 0.5 m_1 1.5 $Y_1(0)$ 1.0 2.0 In terms of masses In terms of cross ratios Area (m)Cross ratios (m)Area (Cross ratios)

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AdS_3 preliminaries

Restrict to simple $R^{1,1}$ kinematics, i.e. surfaces in AdS_3

$$x_{\overline{i}}$$
 $x_{\overline{i+1}}$ $x_{\overline{i+2}}$ $x_{\overline{i}}^+$ $x_{\overline{i+1}}^+$

Flat SL(2) connection $\mathcal{A}(\zeta) = A + \frac{1}{\zeta} \Phi_z dz + \zeta \Phi_{\bar{z}} dz$ $[D_z, D_{\bar{z}}] = 0$ $\Rightarrow \quad p(z) = \operatorname{Tr} \Phi_z^2/2$ Holomorphic function (T(z)=0) Z_2 projection $U\mathcal{A}(\zeta)U^{-1} = \mathcal{A}(-\zeta)$ The area $Area = 2 \int d^2 z \operatorname{Tr}(\Phi_z \Phi_{\bar{z}})$

Boundary conditions

- Simple $|z| \rightarrow \infty$ asymptotics as in four cusp solution
- Having 2n cusps

$$A \sim 0 \quad \text{at large } |z|$$
$$p = z^{n-2} + a_1 z^{n-3} + \dots$$

AdS₃ preliminaries - The linear problem

Flat section
$$[d + \mathcal{A}(\zeta)]\varphi_a = 0$$
, $a = 1, 2$

 $\begin{vmatrix} Y_{-1} + Y_2 & Y_1 - Y_0 \\ Y_1 + Y_0 & Y_{-1} - Y_2 \end{vmatrix} = -1$ Spinor \rightarrow vector problems $Y_{a\dot{a}} = \varphi(1) \cdot M \cdot \varphi_{\dot{a}}(i)$

SL(2) (traceless) connection $\Rightarrow \langle \varphi, \psi \rangle \equiv \epsilon^{\alpha\beta} \varphi_{\alpha} \psi_{\beta} = \text{const}$ (Wronskian)

At
$$|z| \to \infty$$
 $d + \mathcal{A}(\zeta) \sim d + \frac{dz}{\zeta} \begin{pmatrix} \sqrt{p} & 0\\ 0 & -\sqrt{p} \end{pmatrix} + \zeta d\bar{z} \begin{pmatrix} \sqrt{\bar{p}} & 0\\ 0 & -\sqrt{\bar{p}} \end{pmatrix}$, $dw = \sqrt{p}dz$

$$w \to z^{\frac{n}{2}} \quad \Rightarrow \quad \varphi \to e^{\pm (z^{\frac{n}{2}}/\zeta + \overline{z}^{\frac{n}{2}}\zeta)} \quad \text{n Stokes sectors}$$



The cross ratios

Black board derivation

Conclusions

 $\mathcal{A} \sim e^{-\frac{\sqrt{\lambda}}{2\pi}Area}$

Since the area is the free energy, this formula looks like we are computing the partition function of the system on a torus, where one of the sides has length proportional to $\sqrt{\lambda}$.

Future directions

- Continuous limit.
- The quantum problem.
- How to introduce a spectral parameter at weak coupling?
- Correlation functions.



Thank you

Generic AdS₅ Kinematics



AdS₅

$$\log (Y_{2,s}) = -m_s \sqrt{2} \cosh(\theta) - K_2 \star \alpha_s - K_1 \star \beta_s$$

$$\log (Y_{1,s}) = -m_s \cosh(\theta) - C_s - \frac{1}{2} K_2 \star \beta_s - K_1 \star \alpha_s - \frac{1}{2} K_3 \star \gamma_s$$

$$\log (Y_{3,s}) = -m_s \cosh(\theta) + C_s - \frac{1}{2} K_2 \star \beta_s - K_1 \star \alpha_s + \frac{1}{2} K_3 \star \gamma_s$$

where

$$\alpha_s \equiv \log \frac{(1+Y_{1,s})(1+Y_{3,s})}{(1+Y_{2,s-1})(1+Y_{2,s+1})} , \ \gamma_s \equiv \log \frac{(1+Y_{1,s-1})(1+Y_{3,s+1})}{(1+Y_{1,s+1})(1+Y_{3,s-1})}$$
$$\beta_s \equiv \log \frac{(1+Y_{2,s})^2}{(1+Y_{1,s-1})(1+Y_{1,s+1})(1+Y_{3,s-1})(1+Y_{3,s+1})},$$

and

$$K_1 \equiv \frac{1}{2\pi \cosh \theta} , \ K_2 = \frac{\sqrt{2} \cosh \theta}{\pi \cosh 2\theta} , \ K_3 = \frac{i}{\pi} \tanh 2\theta .$$

Finally

$$A_{free} = \sum_{s} \int \frac{d\theta}{2\pi} |m_{s}| \cosh \theta \log \left[(1+Y_{1,s})(1+Y_{3,s})(1+Y_{2,s})^{\sqrt{2}} \right] (\theta + i\alpha_{s})$$



Spacetime cross ratios:

$$U_s^{[r]} \equiv 1 + \frac{1}{Y_{2,s}} \bigg|_{\theta = i\pi r/4}$$

Then

$$\mathcal{U}_{2k-2}^{[0]} = \frac{\mathbf{x}_{-k,k}^2 \mathbf{x}_{-k-1,k-1}^2}{\mathbf{x}_{-k-1,k}^2 \mathbf{x}_{-k,k-1}^2} \; .$$