



# Flux Attractors and Generating Functions

Adapting Black Hole Attractors to String Phenomenology

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# Introduction

- Realistic string theory compactifications must stabilize moduli, e.g. by adding fluxes in the compactification.
- One aspect: *physical variables* (e.g. moduli) are determined as function of *vacuum data* (e.g. fluxes).
- Reminder: the *black hole attractor* mechanism determines the *scalar field at the horizon* in terms of *black hole charges*.
- Goal in this talk: develop *similarities* and *differences* of black hole attractors and *flux attractors*.
- The results will be illuminating *conceptually* and also (somewhat) *practical*: they solve explicit examples.

References: FL and O'Connell: 0905.2130; FL, Robbins, O'Connell: 0912.4448; Anguelova, FL, O'Connell: 1006.4981 .

# Setting: Black Hole Attractors

- Type IIB  $N=2$  compactifications with a prepotential (no global issues).
- Just ***BPS black holes***.
- Just fixed point behavior: no flows, multicenter solutions.
- Focus on  $N = 2$  ***vector fields*** interpreted as D3 branes: decouple hypers.
- ***Supergravity limit***: large black hole charges (no quantization conditions), canonical Kähler potential.

# Setting: Flux Attractors

- Type IIB O3 (some O7?) flux compactifications with a prepotential (no global issues).
- Just *no scale vacua* (these are analogous to BPS).
- Just *fixed point* behavior: no cosmological evolution.
- Focus on *orientifold projections* of  $N = 2$  vector fields but also the inevitable *complex dilaton*.
- *Supergravity limit*: large black hole charges (no quantization conditions, *no tadpole condition*), canonical Kähler potential.

# Black Hole Attractor: Geometry

The spacetime central charge

$$W_{\text{BH}} = \int F_3 \wedge \Omega_3 .$$

Condition of vanishing Kähler derivative

$$D_i W_{\text{BH}} = \int F_3 \wedge D_i \Omega_3 = \int F_3 \wedge \chi_i = 0 .$$

Geometric interpretation: the  $\chi_i$  are  $(2, 1)$  forms so  $F_3$  has no  $(1, 2)$  components.

The condition that  $F_3$  is real then gives the ***geometric form of the attractor equation***

$$F_3 = 2\text{Re}(C\Omega_3) .$$

$C$  is a complex constant with Kähler weight  $(0, 1)$ .

# Flux Attractor: Geometry

The **complex** three-form flux of type IIB on a CY orientifold

$$G_3 = F_3 - \tau H_3 ,$$

gives the low energy  $N = 1$  GVW superpotential

$$W = \int_{\text{CY}} G_3 \wedge \Omega_3 .$$

The  $F_i = 0$  condition is equivalent to  $G_3$  having no  $(1, 2)$  component:

$$F_i = D_i W = \int_{\text{CY}} G_3 \wedge D_i \Omega_3 = \int_{\text{CY}} G_3 \wedge \chi_i .$$

The  $F_\tau = 0$  condition is equivalent to  $G_3$  having no  $(3, 0)$  component:

$$F_\tau = D_\tau W = D_\tau \int_{\text{CY}} G_3 \wedge \Omega_3 = -\frac{1}{\tau - \bar{\tau}} \int_{\text{CY}} \bar{G}_3 \wedge \Omega_3 .$$

**Flux attractor equation in geometric form** (condition that  $G_3$  is  $(0, 3) + (2, 1)$ ):

$$G_3 = \overline{C}\Omega_3 + C^i D_i \Omega_3 .$$

$C, C^i$  are complex constants with Kähler weight  $(0, 1)$ .

Symplectic covariant formulation takes  $C^i D_i \rightarrow L^I D_I$  so

$$G_3 = \overline{C}\Omega_3 + L^I D_I \Omega_3 .$$

The  $L^I$  are subject to the constraint

$$L^I \partial_I K = 0 .$$

The flux attractor equation is equivalent to the familiar ISD (Imaginary-Self-Duality) condition on the complex flux  $G_3$ :

$$*_6 G_3 = i G_3 .$$

Drawback: the moduli are implicit in the ISD form .

# No Scale vs. SUSY

The superpotential does not depend on Kähler moduli so those are not stabilized. Their F-terms are

$$F_a = D_a W = W \partial_a K_t ,$$

with homogeneity ensuring  $\sum_a |\partial_a K_t|^2 = 3$  such that

$$V = e^K \left( \sum_{A=i,\tau,a} |D_A W|^2 - 3|W|^2 \right) = e^K \left( \sum_i |D_i W|^2 + |D_\tau W|^2 \right) .$$

Solutions to  $F_i, F_\tau$ -flatness conditions  $D_i W = D_\tau W = 0$  are absolute minima of the potential and correspond to Minkowski vacua.

These “no-scale” vacua are supersymmetric exactly when  $W = 0$ : if  $W \neq 0$ , the  $F_a \neq 0$  break SUSY.



# Real Basis

Component form of attractor equations: expand in canonical basis of **real** 3-forms  $\{\alpha_I, \beta^I\}$ .

Moduli enter through the complex three-form

$$\Omega_3 = Z^I \alpha_I - F_I \beta^I .$$

The symplectic section  $(Z^I, F_I)$ : a projective coordinate  $Z^I$  on moduli space and also  $F_I = \frac{\partial F}{\partial Z^I}$  ( $F$ =prepotential).

Black hole charges in real basis:

$$F_3 = P^I \alpha_I - Q_I \beta^I .$$

**Complex** fluxes in real basis:

$$G_3 = m^I \alpha_I - e_I \beta^I .$$

# BH Attractors in Components

Black hole attractor equations in component form

$$\begin{aligned}P^I &= 2\text{Re}[CZ^I], \\Q_I &= 2\text{Re}[CF_I].\end{aligned}$$

- Input:  $2(n + 1)$  **real charges**  $\{Q_I, P^I\}$ ,  $I = 0, \dots, n$ .
- Output:  $n$  **complex moduli**, eg.  $z^I = Z^I/Z^0$ ,  $I = 1, \dots, n$ .
- **More output**: the **overall scale**, eg.  $CZ^0$ . It is equivalent to the Kähler invariant spacetime central charge  $CW_{\text{BH}}$ .

# Flux Attractors in Components

Flux attractor equations in component form

$$\begin{aligned} m^I &= \overline{CZ}^I + L^I, \\ e_I &= \overline{CF}_I + L^J F_{JI}. \end{aligned}$$

- Input:  $2(n + 1)$  **complex fluxes**  $\{e_I, m^I\}$ ,  $I = 0, \dots, n$ .
- Output:  $n + 1$  **complex moduli**, eg.  $z^I = Z^I / Z^0$  **and**  $\tau$ .
- More output:  $n + 1$  **complex parameters**  $L^I$  subject to the constraint (redundancy of  $L^I$ ):

$$\overline{CF}_I L^I - \overline{CZ}^I L^J F_{IJ} = 0.$$

- Even more output: the **overall scale**, eg.  $CZ^0$ . It is equivalent to the Kähler invariant spacetime **on-shell superpotential**  $CW$ .

# What do the $L^I$ Mean?

Vanishing first derivatives of the superpotential project out  $(1, 2)$  and  $(3, 0)$  components of the superpotential.

But ***second derivatives of the superpotential are independent parameters***:

$$m_{\alpha\beta} = e^{K/2} D_\alpha D_\beta W .$$

The counting problem revisited: there are  $2(n + 1)$  input fluxes.

The output:  $n + 1$  VEVs of moduli ***and***  $n + 1$  mass parameters.

***The mass parameters can be tuned independently***: they are *not* determined by the VEVs of the moduli.

# Solving BH Attractor Eq's

BH attractor equations:

$$P^I = 2\text{Re}[CZ^I], \quad Q_I = 2\text{Re}[CF_I].$$

Solution in terms of unknown real functions  $\mathcal{S}^I, \mathcal{S}_I$ :

$$2CZ^I = P^I + i\mathcal{S}^I, \quad 2CF_I = Q_I - i\mathcal{S}_I,$$

that **satisfy the Maxwell relations**

$$\frac{\partial \mathcal{S}^I}{\partial P^J} = \frac{\partial \mathcal{S}_J}{\partial Q_I}.$$

So a **generating function** exists such that

$$\mathcal{S}^I = \frac{1}{\pi} \frac{\partial \mathcal{S}}{\partial Q_I}, \quad \mathcal{S}_I = \frac{1}{\pi} \frac{\partial \mathcal{S}}{\partial P^I}.$$

The generating function  $\mathcal{S}$  encodes all dependence of moduli on black hole charges.

# Derivation of Maxwell Relations

$F_I$  are generated by a prepotential so

$$F_I = \partial_I F = \frac{\partial F}{\partial Z^I} .$$

The attractor equations give the exact differential

$$4C^2 dF = (Q_I - i\mathcal{S}_I)(dP^I + id\mathcal{S}^I) .$$

So

$$d(\text{Im}4C^2 F) = -\mathcal{S}_I dP^I + Q_I d\mathcal{S}^I .$$

Legendre transform gives the Maxwell relation

$$d(Q_I \mathcal{S}^I - \text{Im}4C^2 F) = \mathcal{S}_I dP^I + \mathcal{S}^I dQ_I .$$

Bonus: the generating function is related to the prepotential

$$\mathcal{S} = Q_I \mathcal{S}^I - \text{Im}4C^2 F .$$

# Solving Flux Attractor Equations

Solution of attractor equations in terms of potentials  $\phi^I, \theta_I$

$$\begin{aligned}\overline{CZ}^I &= \frac{1}{2}(m^I + \phi^I), & L^I &= \frac{1}{2}(m^I - \phi^I), \\ \overline{CF}_I &= \frac{1}{2}(e_I + \theta_I), & L^J F_{JI} &= \frac{1}{2}(e_I - \theta_I).\end{aligned}$$

The potentials satisfy **Maxwell relations**

$$\frac{\partial \phi^I}{\partial \overline{m}^J} = -\frac{\partial \theta_J}{\partial \overline{e}_I}.$$

So a real **generating function** exists such that

$$\phi^I = 2i\tau_2 \frac{\partial \mathcal{G}}{\partial \overline{e}_I}, \quad \theta_I = -2i\tau_2 \frac{\partial \mathcal{G}}{\partial \overline{m}^I}.$$

$\mathcal{G} = \mathcal{G}(e_I, m^I, \overline{e}_I, \overline{m}^I)$  encodes all the dependence of moduli on fluxes.

# The Complex Dilaton $\tau$

The identification of the generating function  $\mathcal{G}$  above ignores:

1. Manipulations of differentials do not allow variations of  $\tau$ .
2. The constraint (redundancy of  $L^I$ ):

$$0 = e_I \phi^I - \theta_I m^I = e_I \frac{\partial \mathcal{G}}{\partial \bar{e}_I} + m^I \frac{\partial \mathcal{G}}{\partial \bar{m}^I} .$$

Pleasant bonus: these issues “cancel”.

Precise statement: extremizing  $\mathcal{G}$  over  $\tau$  yields the constraint!

There is a practical flexibility: can insert the physical value of  $\tau$  from the outset **or** solve the constraint in the end.



# BH Example: STU-model

$N = 2$  SUGRA model with prepotential:

$$F = \frac{Z^1 Z^2 Z^3}{Z^0} .$$

Geometry: type IIB on orbifold  $T^6/Z^2 \times Z^2 \simeq T^2 \times T^2 \times T^2$ .

The black hole attractor has **8 real charges**: D3 branes with one direction on each  $T^2$  give  $2^3 = 8$  distinct 3-cycles.

The black hole attractor determines **3 complex moduli** and the **complex spacetime central charge**, a total of 8 real parameters.

The generating function is the black hole entropy  $S = \pi \sqrt{I_4}$  where

$$I_4 = 4(Q_0 P^1 P^2 P^3 - P^0 Q_1 Q_2 Q_3) - \sum_I (P^I Q_I)^2 - 2P^0 Q_0 \sum_i P^i Q_i + 2 \sum_{i < j} P^i Q_j P^j Q_i .$$

# Flux Example: STU-model

Geometry: type *IIB orientifold* on  $T^6/Z^2 \times Z^2 \simeq T^2 \times T^2 \times T^2$ .

Input parameters: **16 real fluxes**. Each of the  $\delta$  three-cycles supports  $H_3$  **and**  $F_3$ .

Primary outputs: **3 complex moduli** and the **complex dilaton**, a total of  $\delta$  real parameters.

Additional outputs: **mass parameters** of the moduli and also the **gravitino mass**.

# Generating Function

Only determined explicitly for **reduced fluxes**, a subset of “just” 8 real fluxes satisfying simple reality conditions:

$$\bar{m}^0 = m^0, \quad \bar{e}_i = e_i, \quad \bar{m}^i = -m^i, \quad \bar{e}_0 = -e_0.$$

The **generating function for reduced fluxes**:

$$\begin{aligned} \mathcal{G} &= \frac{1}{2} \left[ -e_0^h m_f^0 + e_i^f m_h^i \right] + \sum_i \operatorname{sgn}(m_f^0 m_h^i) \sqrt{-e_0^h m_f^0} \sqrt{m_h^i e_i^f} \\ &\quad - \sum_{i < j} \operatorname{sgn}(m_h^i m_h^j) \sqrt{m_h^i e_i^f} \sqrt{m_h^j e_j^f}. \end{aligned}$$

No obvious relation to the black hole entropy. (No “analytic continuation”, it seems).

It is interesting to find the generating function for more general fluxes (not just reduced fluxes), other compactifications. . . .

# The Explicit Moduli

Coupling constant:

$$\text{Im}\tau = \left( -\frac{m_f^0 e_1^f e_2^f e_3^f}{e_0^h m_h^1 m_h^2 m_h^3} \right)^{1/4} .$$

Moduli:

$$z^1 = -i \left[ \left( -\frac{e_0^h}{m_f^0} \right) \left( \frac{m_h^1}{e_1^f} \right) \left( \frac{e_2^f}{m_h^2} \right) \left( \frac{e_3^f}{m_h^3} \right) \right]^{1/4} ,$$
$$z^2 = -i \left[ \left( -\frac{e_0^h}{m_f^0} \right) \left( \frac{e_1^f}{m_h^1} \right) \left( \frac{m_h^2}{e_2^f} \right) \left( \frac{e_3^f}{m_h^3} \right) \right]^{1/4} ,$$
$$z^3 = -i \left[ \left( -\frac{e_0^h}{m_f^0} \right) \left( \frac{e_1^f}{m_h^1} \right) \left( \frac{e_2^f}{m_h^2} \right) \left( \frac{m_h^3}{e_3^f} \right) \right]^{1/4} .$$

# Sign Restrictions

No-scale solutions exist only when the fluxes satisfy certain conditions.

Condition that ***complex structures are in the upper half-plane*** (coupling constants are positive) imposes

$$-\text{sgn}(m_f^0 e_0^h) = \text{sgn}(m_h^1 e_1^f) = \text{sgn}(m_h^2 e_2^f) = \text{sgn}(m_h^3 e_3^f) = +1 .$$

If these are violated: the superpotential is minimized in the unphysical region ( $T^2$  volumes are not positive).

If the ***product*** of these conditions is not satisfied there are not even unphysical solutions: this is the analogue of the non-BPS branch of black holes.

Such input fluxes do not correspond to solutions with Minkowski geometry — they constitute the ***de Sitter branch***.

# Restoring General Fluxes

The explicit STU-solution involves just 8 fluxes and (therefore) purely imaginary moduli.

The model has large symmetry:  $SL(2)^4$

This symmetry generalizes the  $SL(2)^3$  of STU black holes to include  $S$ -duality acting on the  $\tau$  modulus.

The formulae for moduli can be *uniquely* generalized to the entire  $SL(2)^4$  orbit, but the details have not been carried out.

Explicit expressions appear complicated because there are 4 invariants (for BHs there is just one, the quartic invariant  $I_4$  that determines the entropy completely).

# Invariants

The simplest invariant is quadratic and very well known:

$$I_2 = \int F_3 \wedge G_3 = -e_0^h m_f^0 + e_i^f m_h^i$$

But there are **4 independent invariants** of  $SL(2)^4$ : the quadratic  $I_2$ , two quartic  $I_4^{1,2}$ , and one sextic  $I_6$ .

The general solution to the STU-example of the flux attractor equations are determined uniquely by these invariants (at least locally).

# Generating Function: Properties

$\mathcal{G}$  is homogeneous of degree  $(1, 1)$ : scaling of  $e_I, m^I$  scales  $\mathcal{G}$  so

$$\mathcal{G} = e_I \frac{\partial \mathcal{G}}{\partial e_I} + m^I \frac{\partial \mathcal{G}}{\partial m^I} = \frac{i}{2\tau_2} (e_I \bar{\phi}^I - m^I \bar{\theta}_I) .$$

The RHS is real (but not manifestly so).

The ***on-shell superpotential*** is related to the generating function

$$-\frac{i}{2\tau_2} CW = \int F_3 \wedge H_3 - \mathcal{G} .$$

The ***on-shell*** Kähler potential is similarly related to fluxes, and so is the ***gravitino mass***

$$m_{3/2}^2 = e^K |W|^2 = -i \overline{CW} = 2\tau_2 \left( \int F_3 \wedge H_3 - \mathcal{G} \right) .$$



# Positivity

The generating function  $\mathcal{G}$  defined in analogy with the entropy function  $\mathcal{S}$  is not positive definite.

The positive function in the flux ensemble is the **gravitino mass** (squared), considered **as function of all fluxes**.

In the STU-example:

$$m_{3/2}^2 = \frac{1}{2} \left( \text{sgn}(m_f^0) \sqrt{-e_0^h m_f^0} + \sum_i \text{sgn}(m_h^i) \sqrt{m_h^i e_i^f} \right)^2 .$$

# Generalizations

- This talk: standard type IIB (GKP) scenario (with Kähler moduli unstabilized).
- Generalization: stabilize Kähler moduli perturbatively using ***Geometric Fluxes***. (with O'Connell and Robbins).
- Consider ***Heterotic Compactifications***. (with Angelova and O'Connell)
- Consider  **$G_2$  *Compactifications of M-theory***. (unpublished, with O'Connell and Robbins)

# Generating Function: Interpretation

The generating function is the flux attractor analogue of the black hole entropy.

Is it (or one of its relatives) related to a counting problem for flux vacua?

A candidate: “counting” of flux vacua usually involves enumeration of distinct flux quantum numbers.

It is natural to expect that ***a given set of semi-classical flux quantum numbers can be realized microscopically in many ways.***

The generating function  $\mathcal{G}$  may count such realizations semi-classically.

The upshot: a thermodynamic interpretation of the generating function may give a short-cut to the ***measure on the landscape.***

# Summary

- ***Studied IIB orientifold compactifications*** with ISD fluxes
- ***Analyzed attractor equations***: the input fluxes determine scalar moduli ***and*** mass parameters.
- Solved attractor equations in terms of a ***Generating Function***, a scalar function of the fluxes.
- Explicit computations for a canonical example with ***16 fluxes***.
- Speculated that the generating function constitutes a ***measure on moduli space***.